Central Bureau for Astronomical Telegrams INTERNATIONAL ASTRONOMICAL UNION CBAT Director: Daniel W. E. Green; Hoffman Lab 209; Harvard University; 20 Oxford St.; Cambridge, MA 02138; U.S.A. e-mail: cbatiau@eps.harvard.edu (alternate cbat@iau.org) URL http://www.cbat.eps.harvard.edu/index.html Prepared using the Tamkin Foundation Computer Network

## (2623) ZECH

D. Pray, Sugarloaf Mountain Observatory, South Deerfield, MA, U.S.A.; P. Pravec, Ondrejov Observatory; V. Chiorny and Yu. N. Krugly, Kharkiv Observatory; V. Benishek, Belgrade Observatory; J. Pollock, D. B. Caton, A. B. Smith, and R. L. Hawkins, Appalachian State University; R. Inasaridze, Abastumani Observatory; R. Montaigut and A. Leroy, OPERA Observatory, France; and D. E. Reichart, J. B. Haislip and A. P. LaCluyze, University of North Carolina, Chapel Hill, report that photometric observations obtained during 2014 Oct. 7-31 reveal that minor planet (2623) is a binary system with an orbital period of 117.2 +/- 0.3 hr. The primary shows a period of 2.7401 +/-0.0002 hr, and it has a lightcurve amplitude of 0.22 mag at solar phases 13-25 deg, suggesting a nearly spheroidal shape. Mutual eclipse/occultation events that are 0.09-magnitude deep indicate a lower limit on the secondary-to-primary mean-diameter ratio of 0.29. The secondary is non-synchronous, and it has a period of 18.718 +/- 0.008 hr and an amplitude of 0.08 mag in the observed lightcurve, suggesting a secondary elongation of about 2:1 after removing contribution of light from the primary.

NOTE: These 'Central Bureau Electronic Telegrams' are sometimes superseded by text appearing later in the printed IAU Circulars.

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